

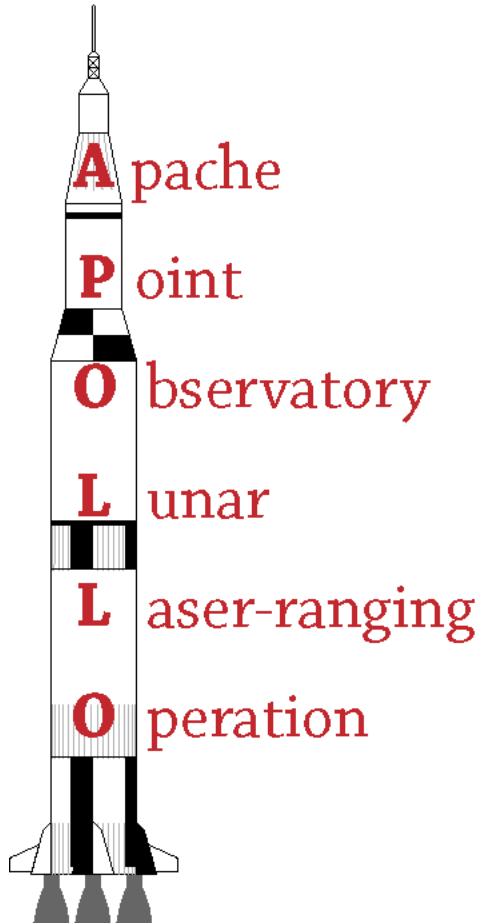
A large, glowing orange-red moon is positioned on the left side of the frame. A bright green laser beam originates from the right edge and points towards the moon. The background is a dark, starry space.

Shining Dark on Dust

Using Total Lunar Eclipses
to expose
Dust Accumulation on Apollo Reflectors

Tom Murphy (UCSD)

APOLLO: one giant leap for LLR



- APOLLO performs lunar laser ranging (LLR) to test the foundations of gravity
 - Is General Relativity (GR) correct?
 - Equivalence Principle violation?
 - Time variation of gravitational strength?
 - Departure from $1/r^2$ force law?
 - Gravitomagnetism (GR effect)
 - Geodetic Precession
- APOLLO gets millimeter range precision
 - GR departures from Newton ~ 10 m level
- Acronym assures funding by NASA
 - $\sim 50/50$ NASA/NSF

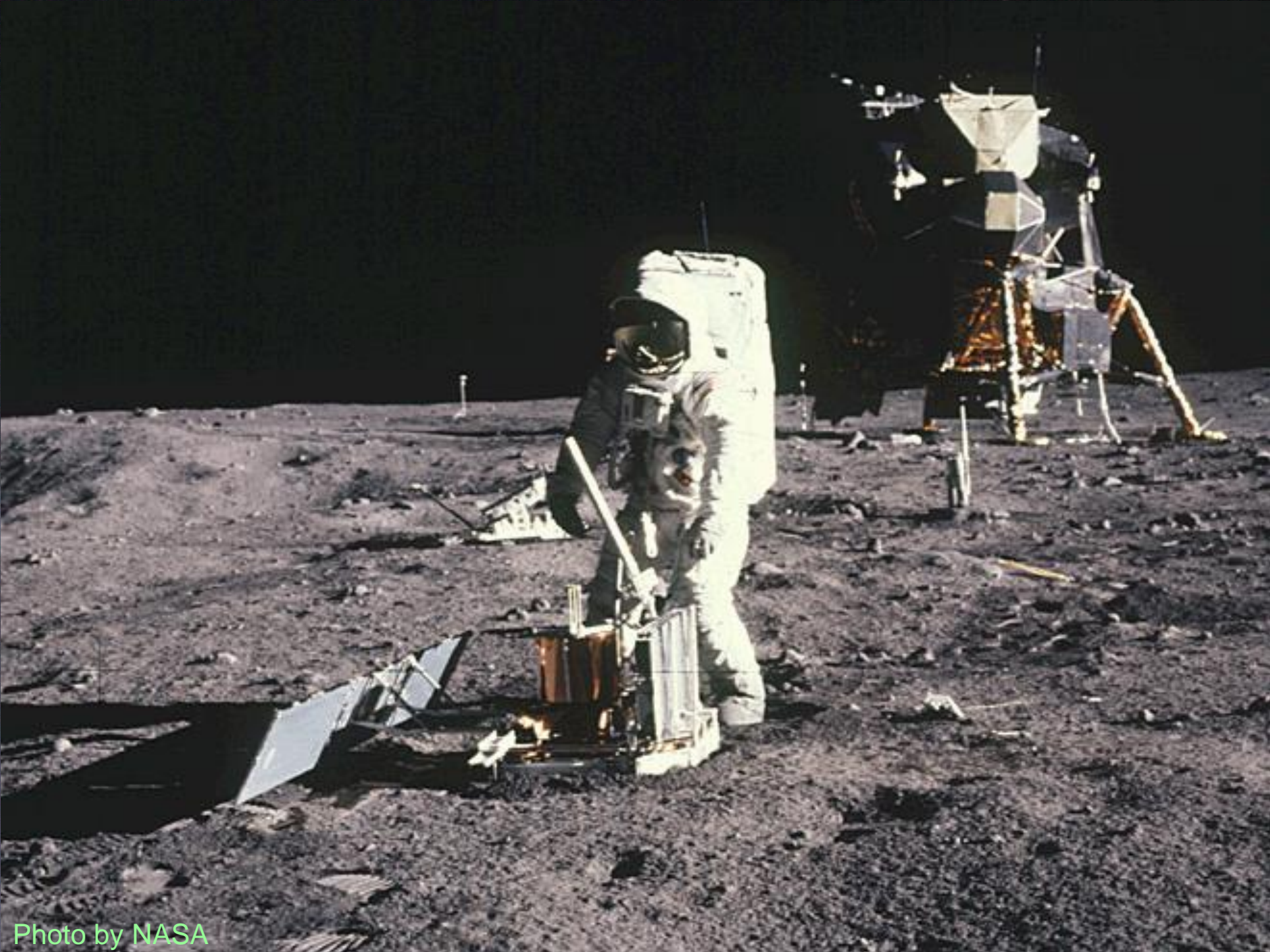
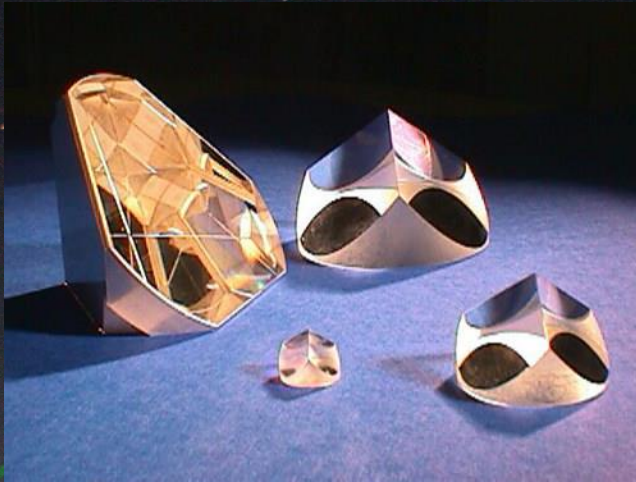


Photo by NASA

Lunar Retroreflector Arrays



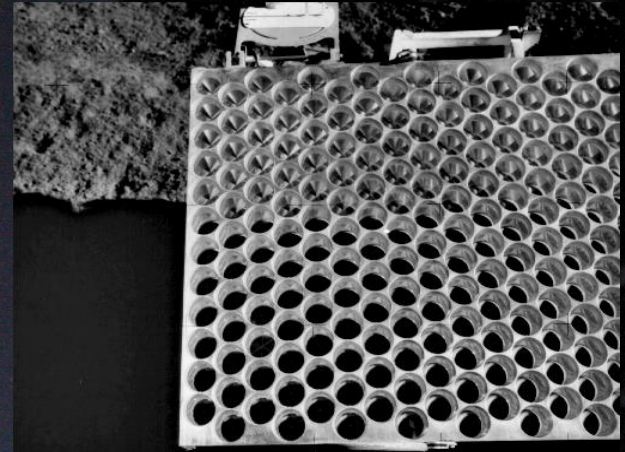
Corner cubes



Apollo 11 retroreflector array



Apollo 14 retroreflector array



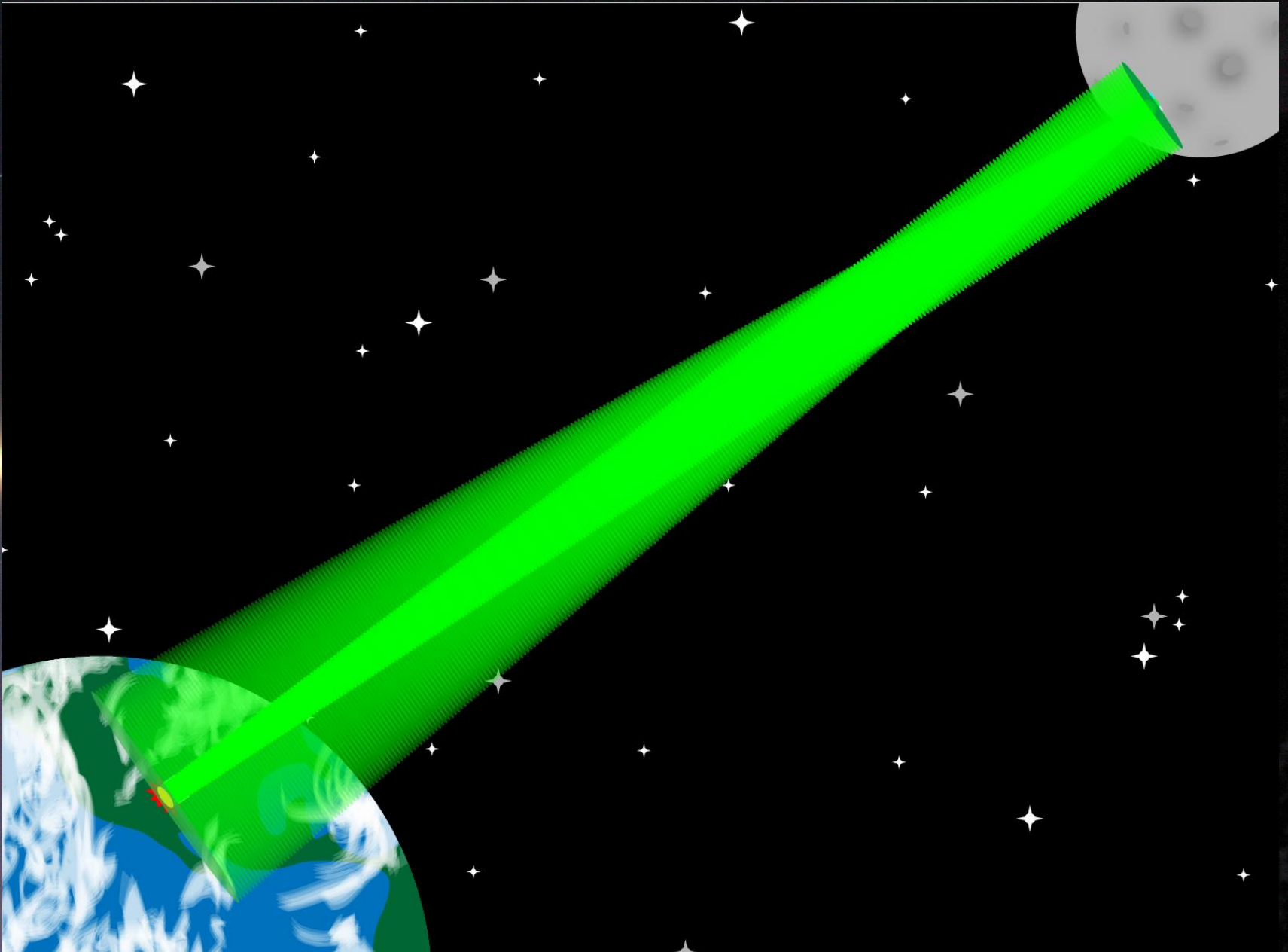
Apollo 15 retroreflector array

The Reflector Positions



- Three Apollo missions left reflectors
 - Apollo 11: 100-element
 - Apollo 14: 100-element
 - Apollo 15: 300-element
- Two French-built, Soviet-landed reflectors were placed on rovers
 - Luna 17: Lunokhod 1 rover
 - Luna 21: Lunokhod 2 rover
 - similar in size to A11, A14
- Signal loss is huge:
 - $\approx 10^{-8}$ of photons launched find reflector (atmospheric seeing)
 - $\approx 10^{-8}$ of returned photons find telescope (corner cube diffraction)
 - $>10^{17}$ loss considering other optical/detection losses

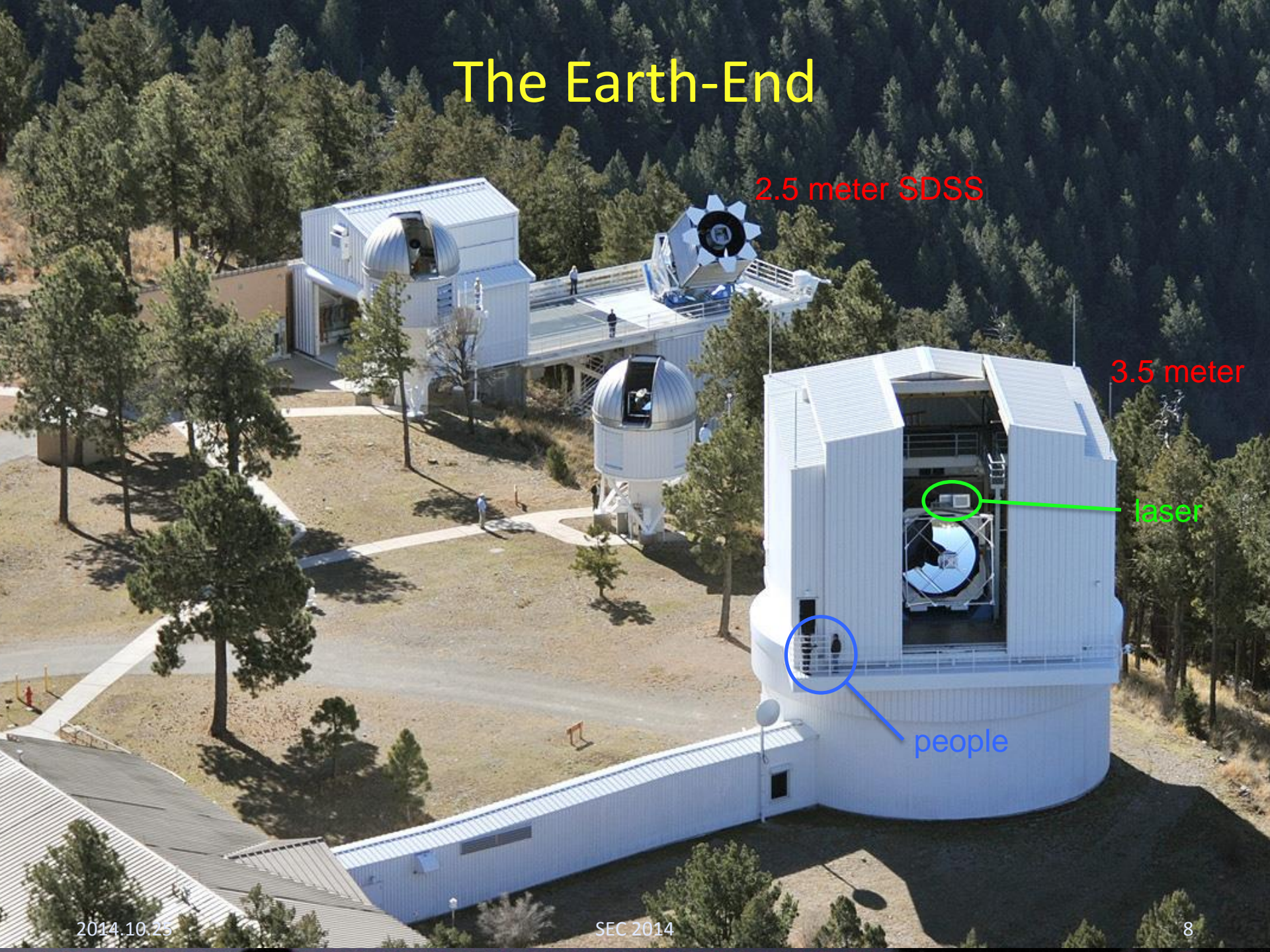
How Does it Work?



Big Bang Theory: Making it Look Easy



The Earth-End



2.5 meter SDSS

3.5 meter

laser

people

APOLLO's Secret Weapon: Aperture



- The Apache Point Observatory's 3.5 meter telescope
 - Southern NM (Sunspot)
 - 9,200 ft (2800 m) elevation
 - Great “seeing”: 1 arcsec
 - Flexibly scheduled, high-class research telescope
 - APOLLO gets 8–10 < 1 hour sessions per lunar month
 - 7-university consortium (UW NMSU, U Chicago, Princeton, Johns Hopkins, Colorado, Virginia)

APOLLO Laser

- Nd:YAG; flashlamp-pumped; mode-locked; cavity-dumped
- Frequency-doubled to 532 nm
 - 57% conversion efficiency
- 90 ps pulse width (FWHM)
- 115 mJ (green) per pulse
 - after double-pass amplifier
- 20 Hz pulse repetition rate
- 2.3 Watt average power
- GW peak power!!
- Beam is expanded to 3.5 meter aperture
 - Less of an eye hazard
 - Less damaging to optics

Laser Mounted on Telescope



A Telescope in Reverse



Gigantic Laser Pointer





2.5 meter Telescope
University of Arizona

photo by Dan Long

Killer Returns

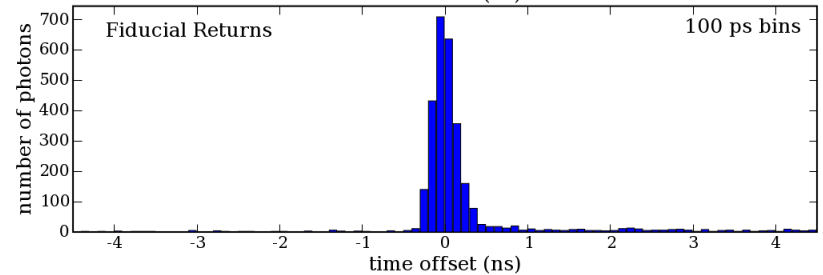
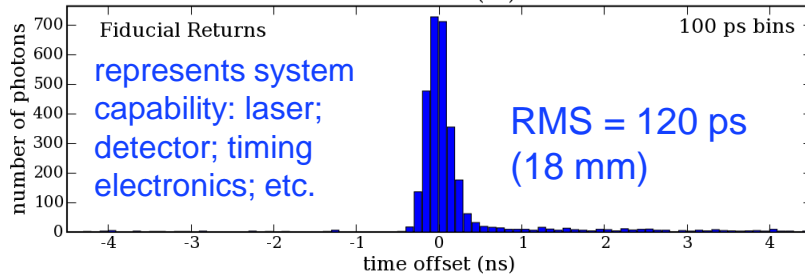
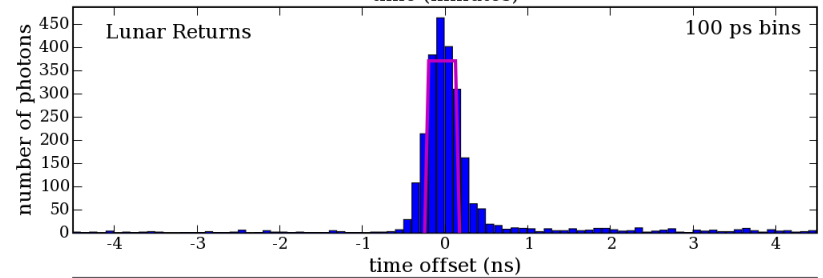
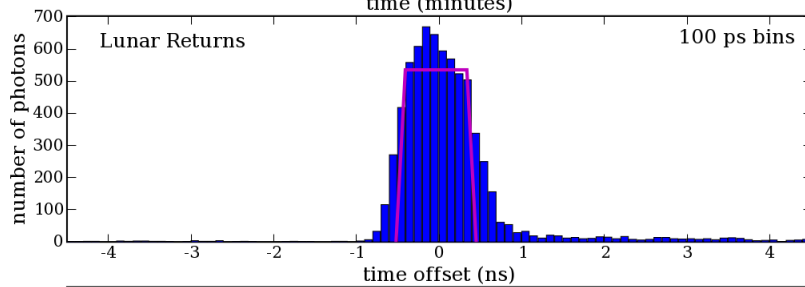
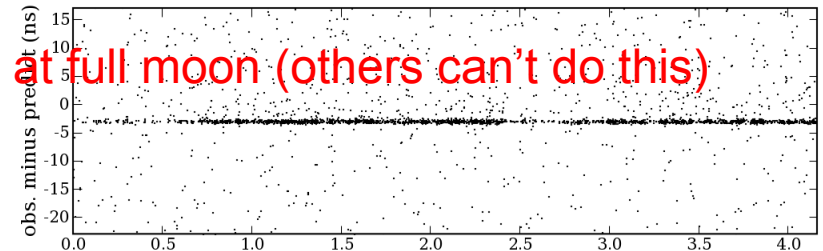
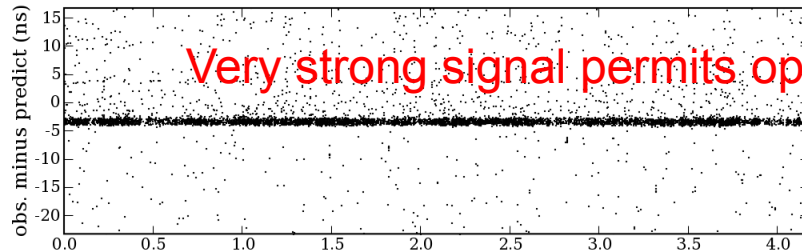
Apollo 15

2007.11.19

Apollo 11

red curves are theoretical profiles: get convolved with fiducial to make lunar return

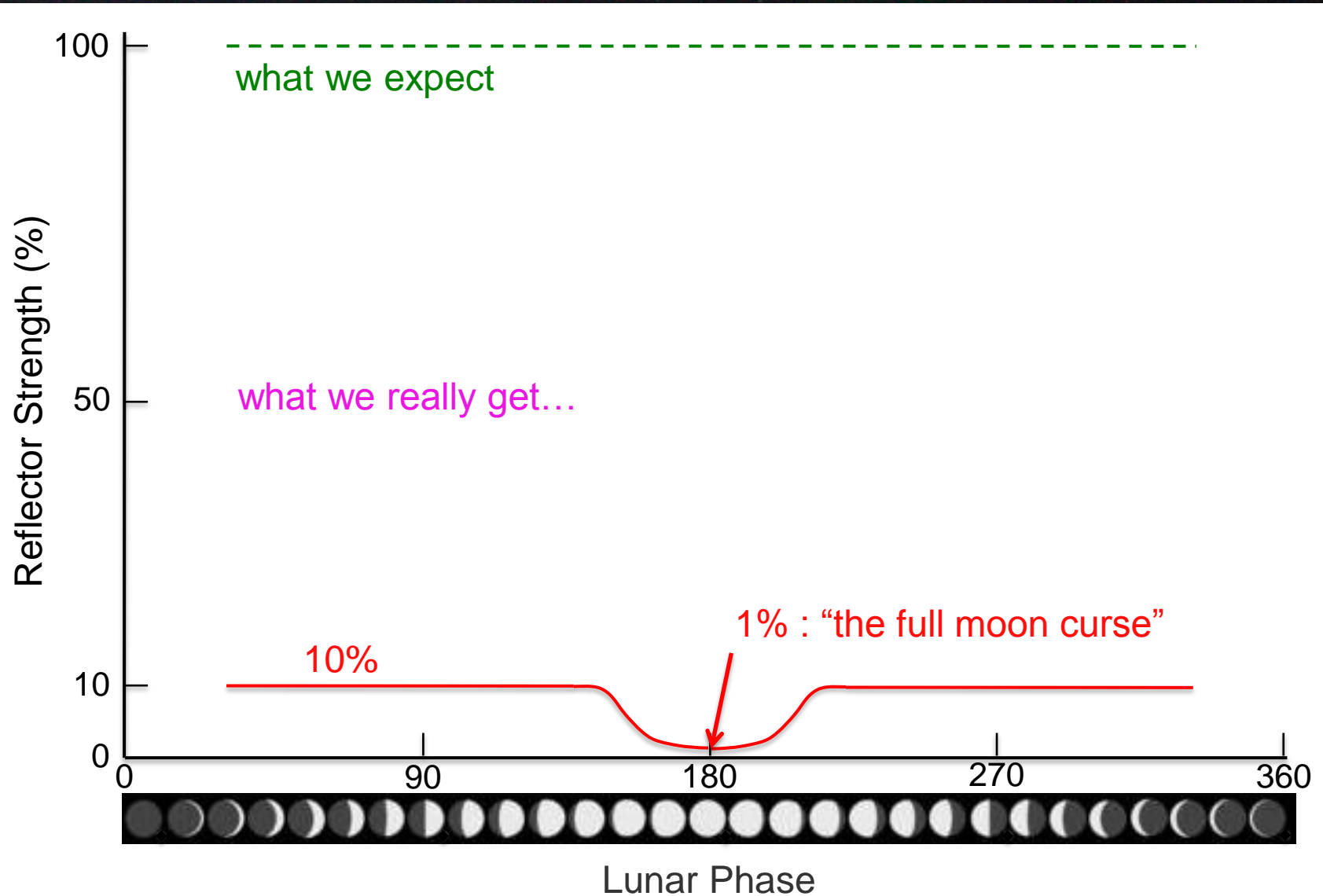
Very strong signal permits operation at full moon (others can't do this)



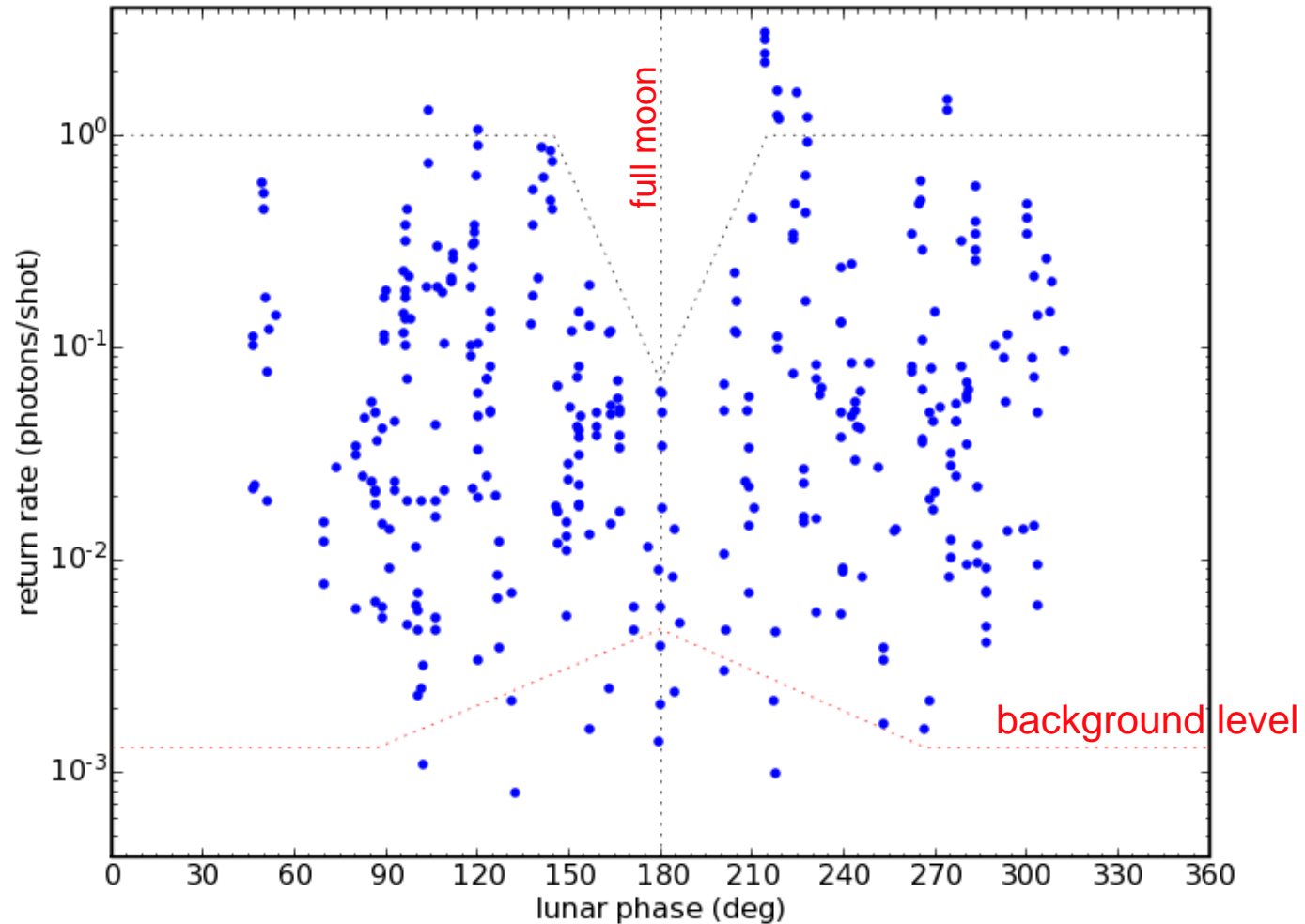
- 6624 photons in 5000 shots
- 369,840,578,287.4 \pm 0.8 mm
- 4 detections with 10 photons

- 2344 photons in 5000 shots
- 369,817,674,951.1 \pm 0.7 mm
- 1 detection with 8 photons

Not All is Rosy in LLR-Land

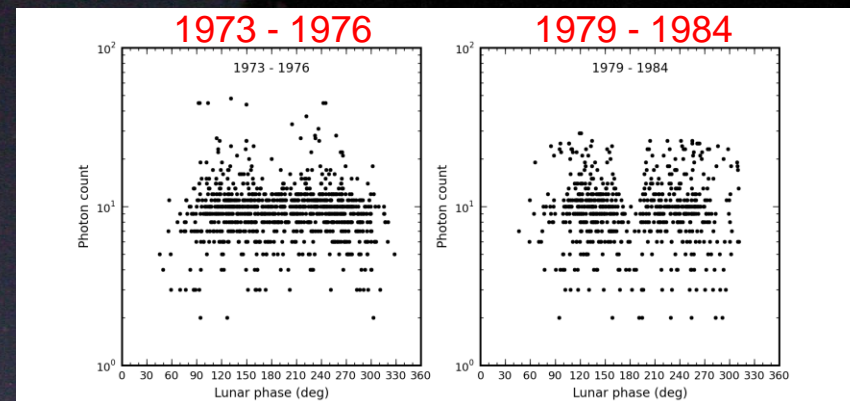
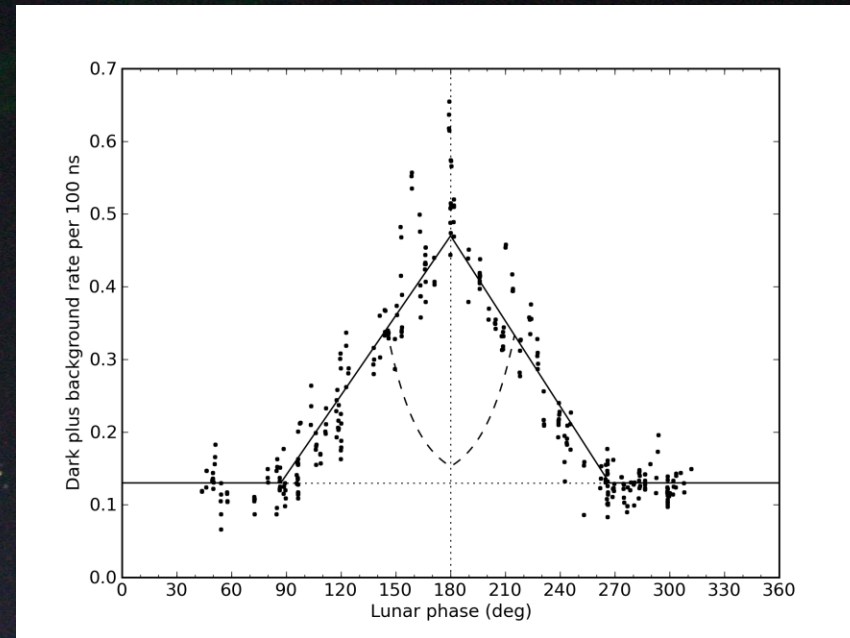


APOLLO rates on Apollo 15 reflector

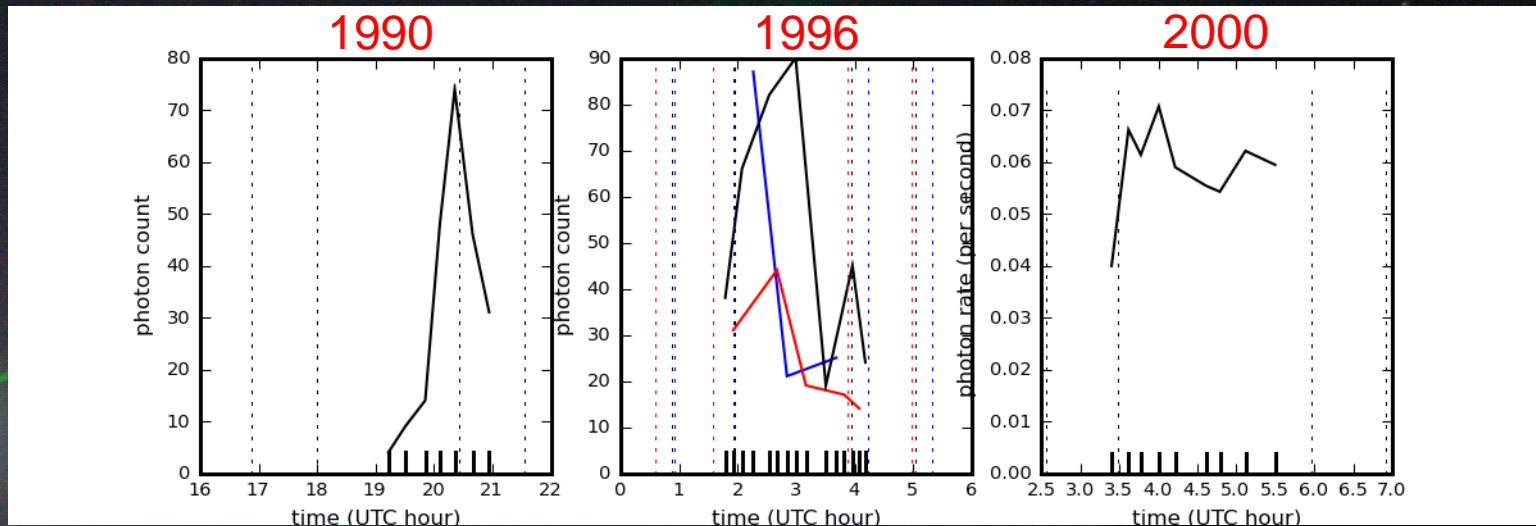


More on the deficit

- APOLLO system sensitivity is not to blame for full-moon deficit
 - background is not impacted
- Early LLR data trucked right through full-moon with no problem
- The deficit **began to appear** around **1979**
- No full-moon ranges from 1985 until 2006, except during **eclipse**

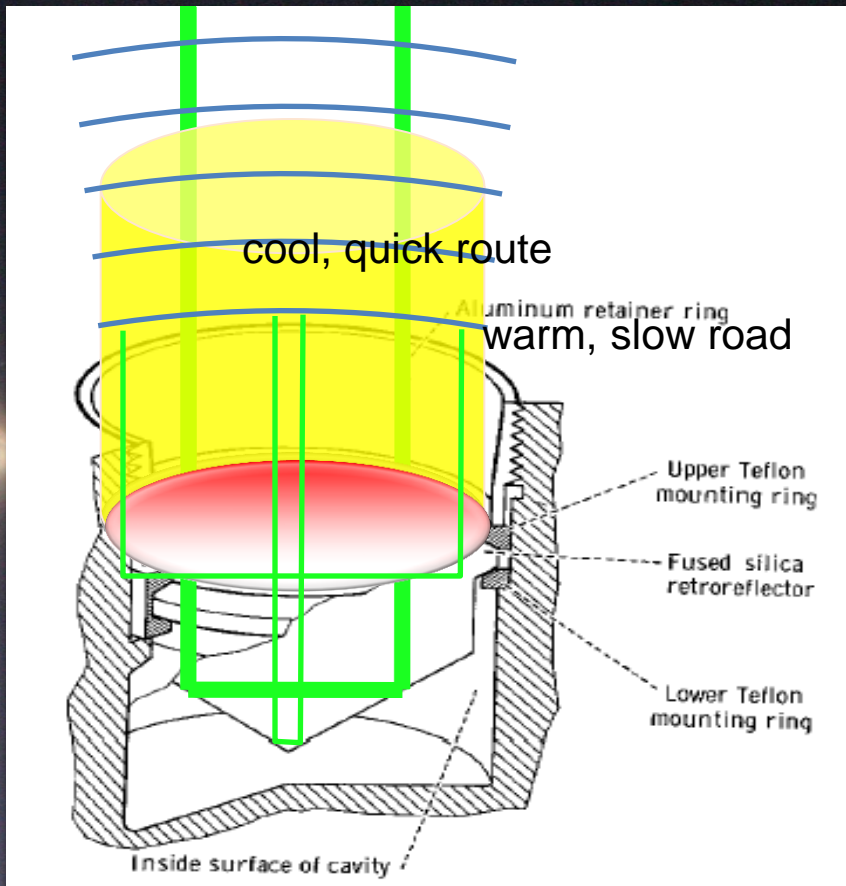


Past Eclipses, French Observations



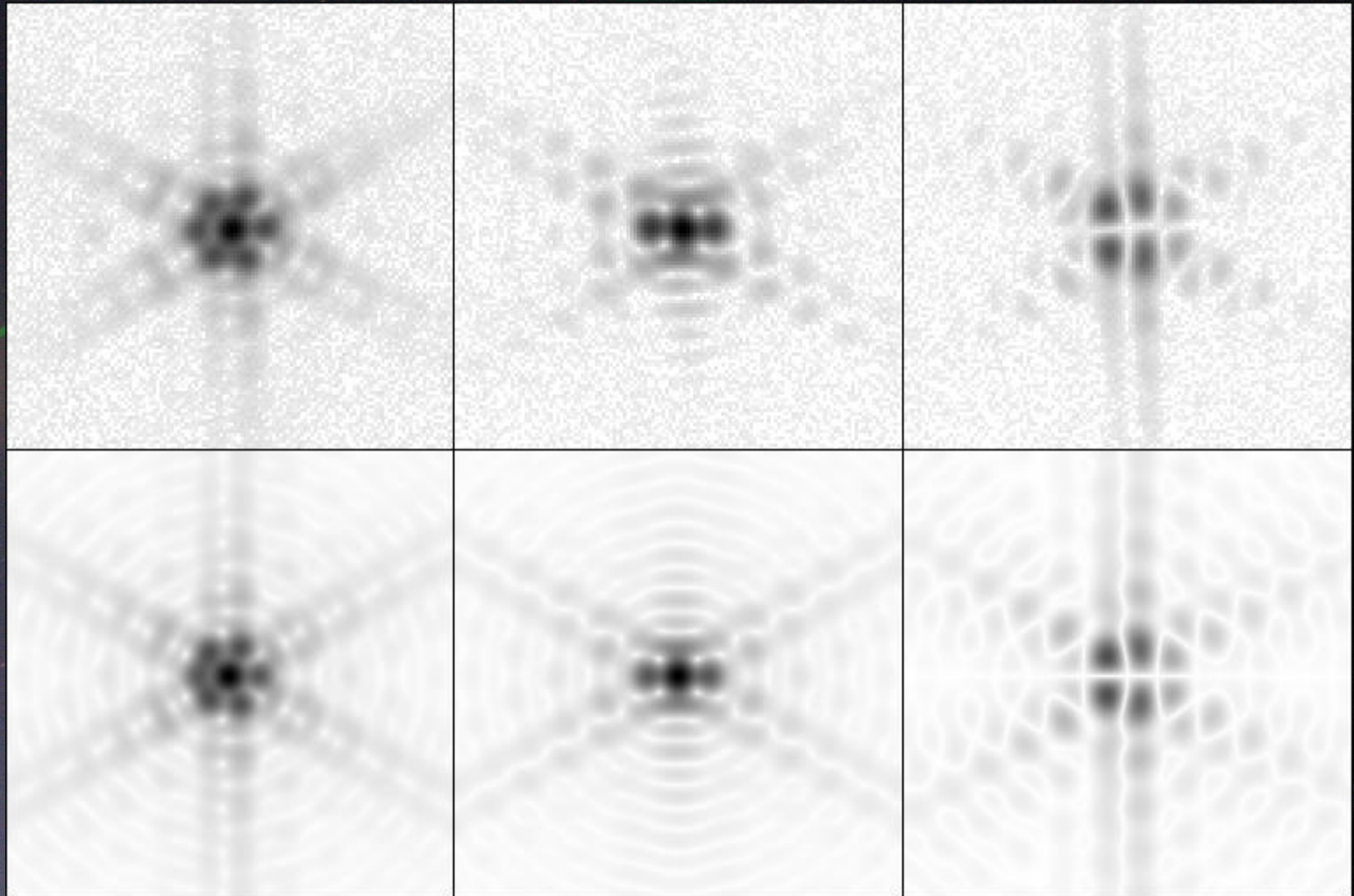
- Strong signal during eclipse
 - Apollo 11 (blue) was about as strong as this station saw in decades of ranging: definitely a special night
- Take your pick: late peak; early peak; no peak
- LLR is hard: ups and downs can be acquisition difficulty

What's Wrong?

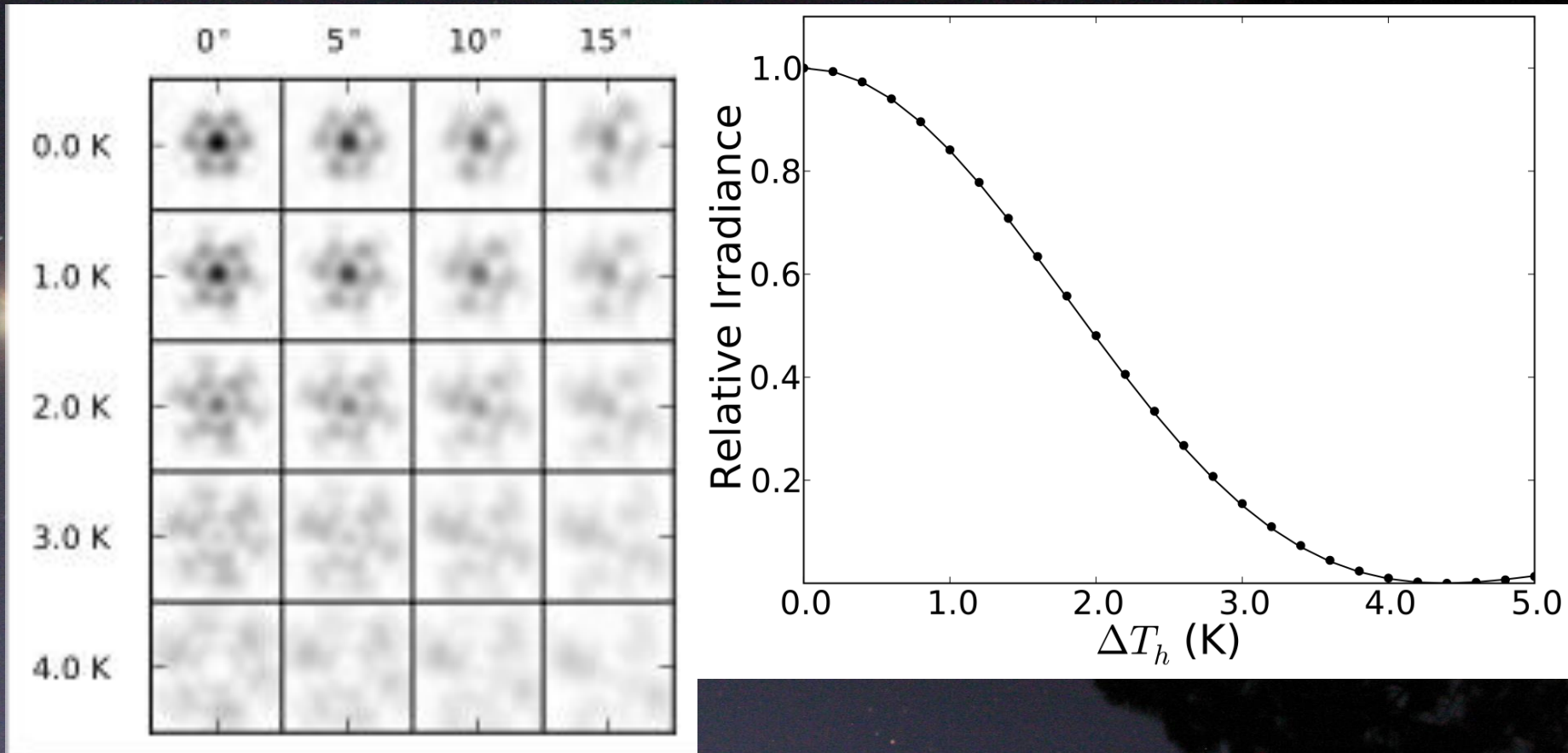


- The **full-moon deficit**, together with **normal eclipse behavior**, gives us the best clues:
 - **thermal** nature
 - absorbing **solar flux**
- Most likely: **dust**
 - Obviously could explain overall deficit (10%)
- Full moon effect then due to solar heating of dust
 - sun comes straight down tube at full moon
 - makes front hotter than vertex of corner cube, leading to divergence of exit beam
 - only takes 4°C (7°F) gradient to introduce $10\times$ reduction

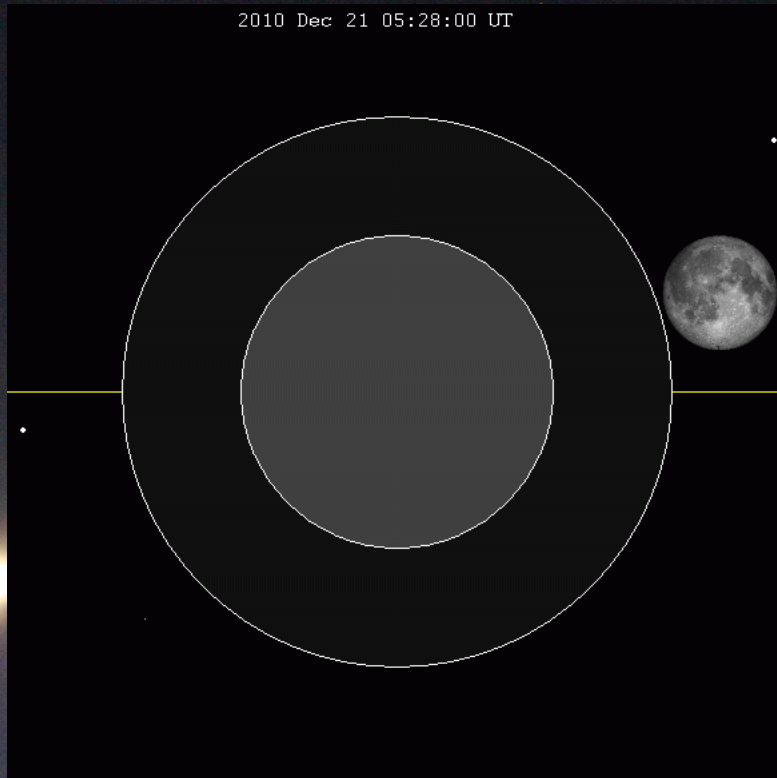
Modeling CCR Diffraction Patterns



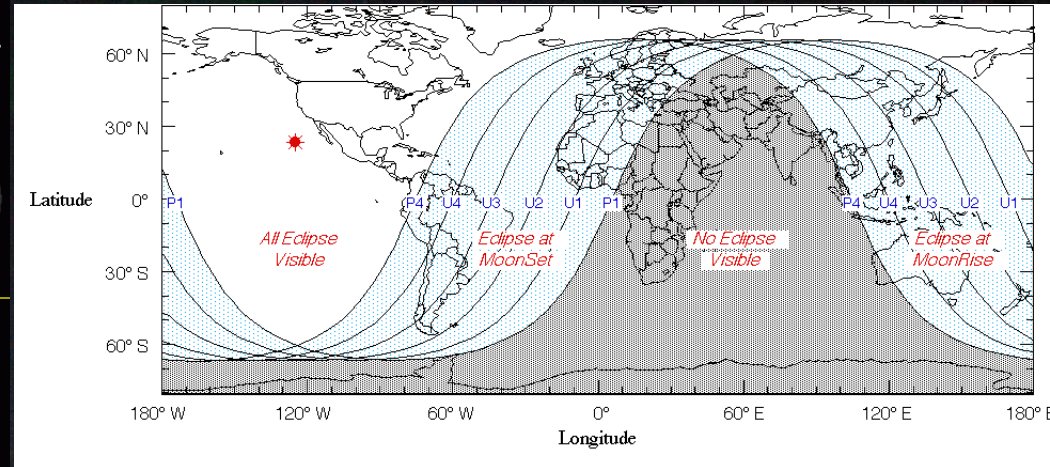
Exploring Orientation & Thermal Gradients



Eclipse as Light Switch

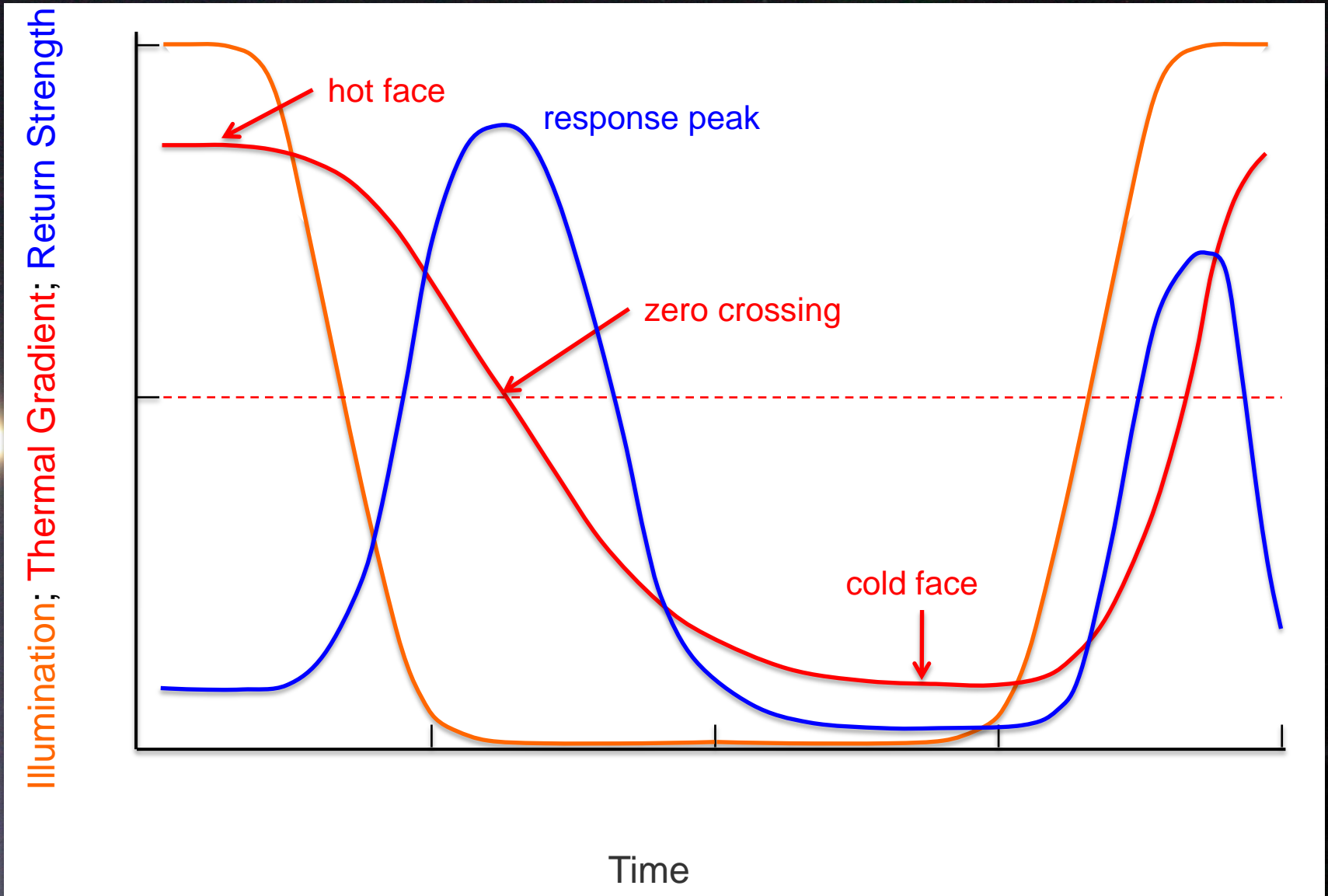


Dec. 2010: perfect eclipse for North America



- If sunlight is to blame, let's shut it off at full moon!
 - need to intercede with massive body: move heaven and earth?
 - examine response time: is it a thermal effect in corner cubes?

Cartoon of Expectations

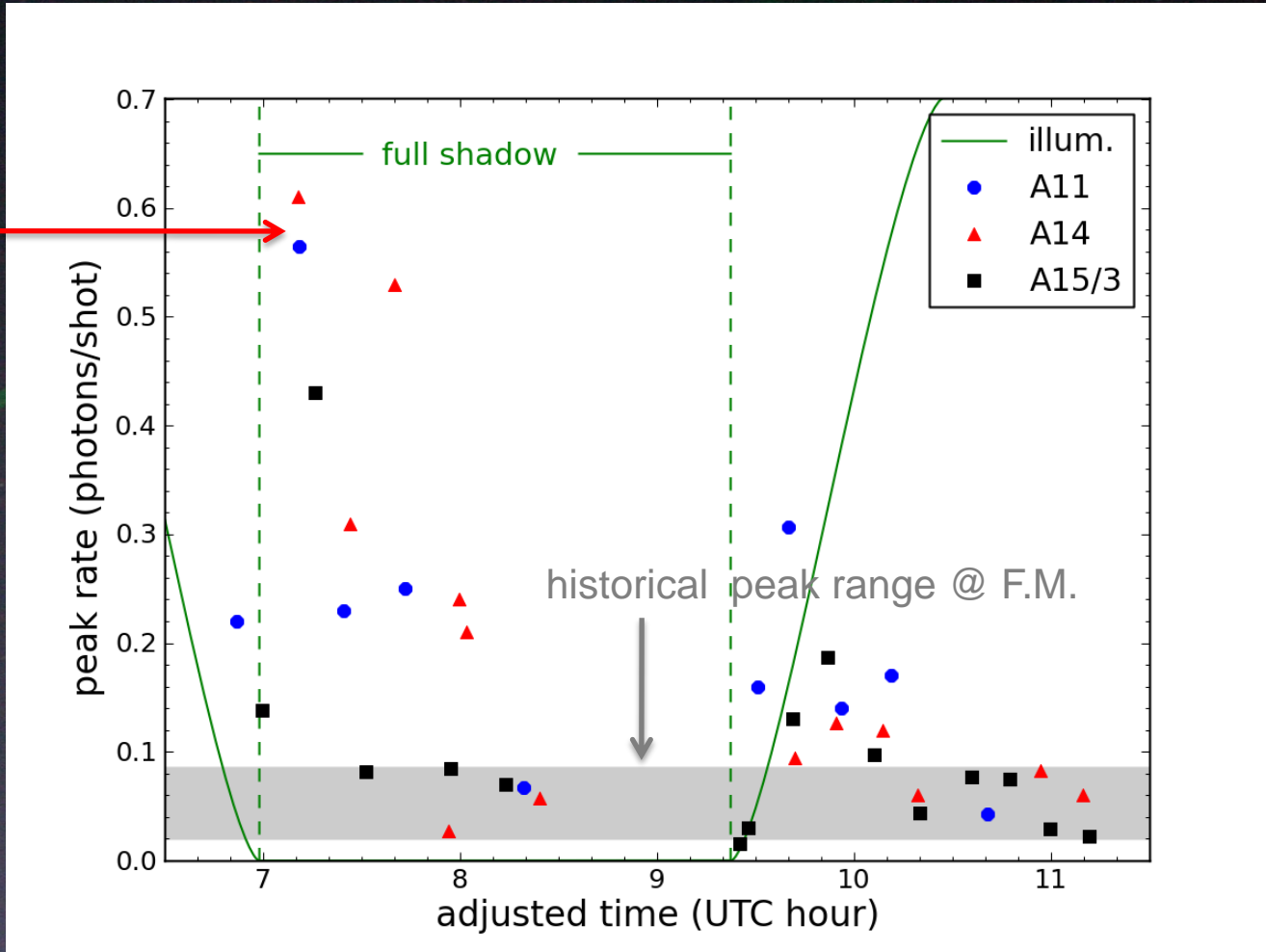


December 21, 2010

Near-zenith at mid eclipse for APOLLO
but variable, high, thin clouds that night



2010 APOLLO Eclipse Results

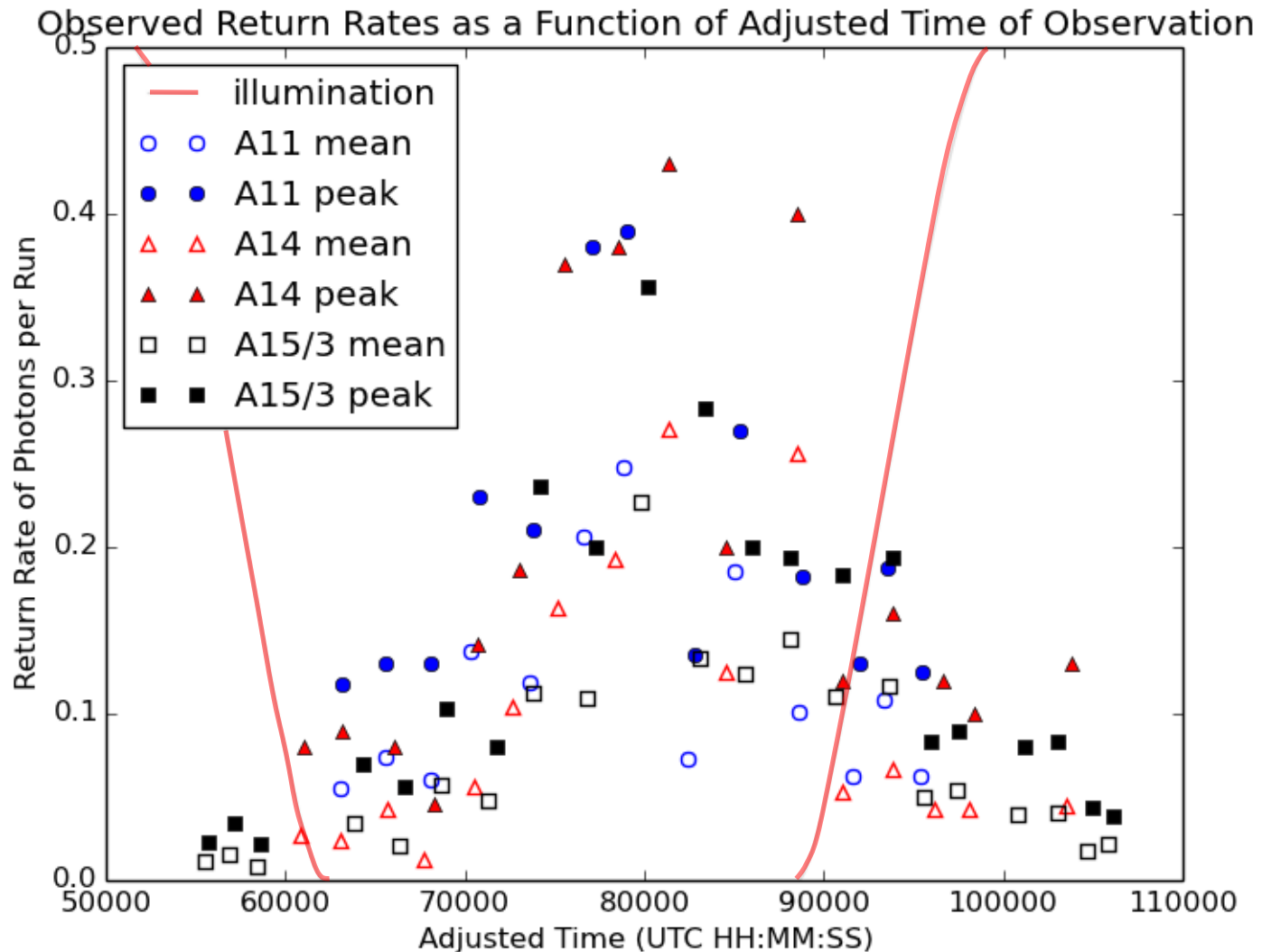


robust recovery initially, then down, and brief resurgence once light returns

2014.04.15 Eclipse



Still See Dramatic Effect, but Single-Peak



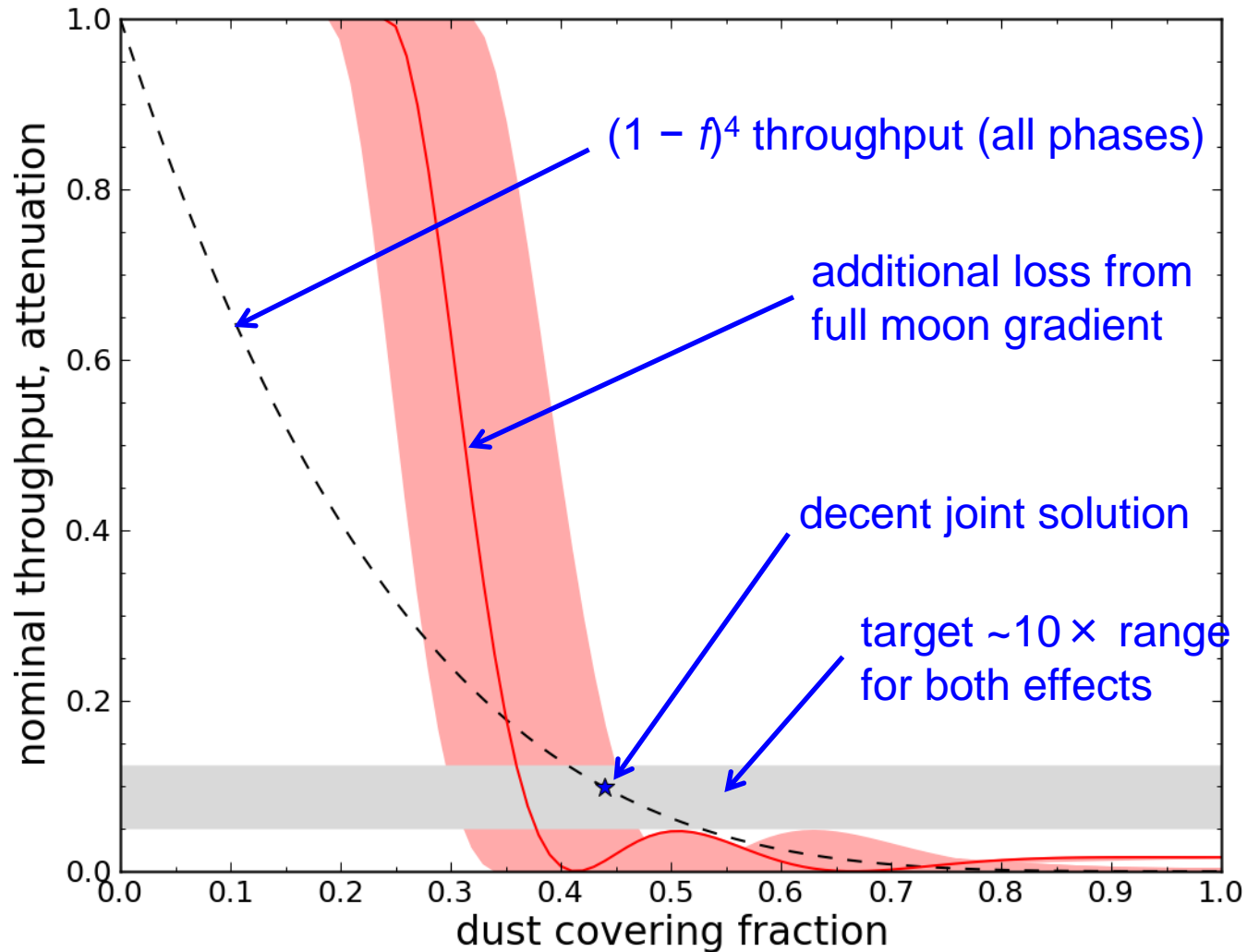
Reconciliation?

- Thicker set of **clouds** during 2010 eclipse coincide with dip
 - analysis of opacity suggested this wasn't responsible
- Conditions in 2014 eclipse **pristine**
 - and signal was acquired well before umbral stage commenced
- Must conclude that **single peak** is correct
 - cartoon predicting double peak got "lucky!"
 - time constant is longer: thermal coupling to aluminum pallet

What CAN We Say?

- Thermal effect real: solar absorption happening
 - likely dust coating
- Roughly $10 \times$ signal loss over expectations, at all phases
- Factor of 10–15 additional signal loss at full moon
 - recovering to admirably strong levels during both eclipses
 - consistent with thermal gradients in 3–4 K range at full moon
- Putting together: $10 \times$ attenuation plus large gradient
 - suggests dust covering fraction is $f \approx 0.4\text{--}0.5\%$
 - double-pass and diffraction result in far-field intensity $(1 - f)^4$
 - similar fraction computed from radiative balance to get gradient

Covering Fraction



Summary

- **APOLLO** is a millimeter-capable lunar ranging station testing gravity
- Strong signal allows LLR operation at full-moon
- Found that reflectors were “sick” near full moon; suspected thermal/solar issue
- Eclipse provides celestial light switch to test idea
- **CONFIRMED**: definite solar/thermal effect, likely due to dust deposition (signal levels outstanding during eclipse)
- **ESTIMATE**: dust covering fraction of nearly 50%
 - roughly a mono-layer per century

